

Seasonal variation of radial brightness contrast of Saturn's rings viewed in mid-infrared by COMICS

Fujiwara et al. 2017, A&A, 599, A29

Hideaki Fujiwara
Subaru Telescope, NAOJ

Collaborators: Ryuji Morishima (UCLA/JPL),
Takuya Fujiyoshi, Takuya Yamashita (NAOJ)

Planetary Rings: Miniature of dusty debris disks?

- Disk structure by solid particles
- Gas-free or gas-less system
- Probably generated by collisions

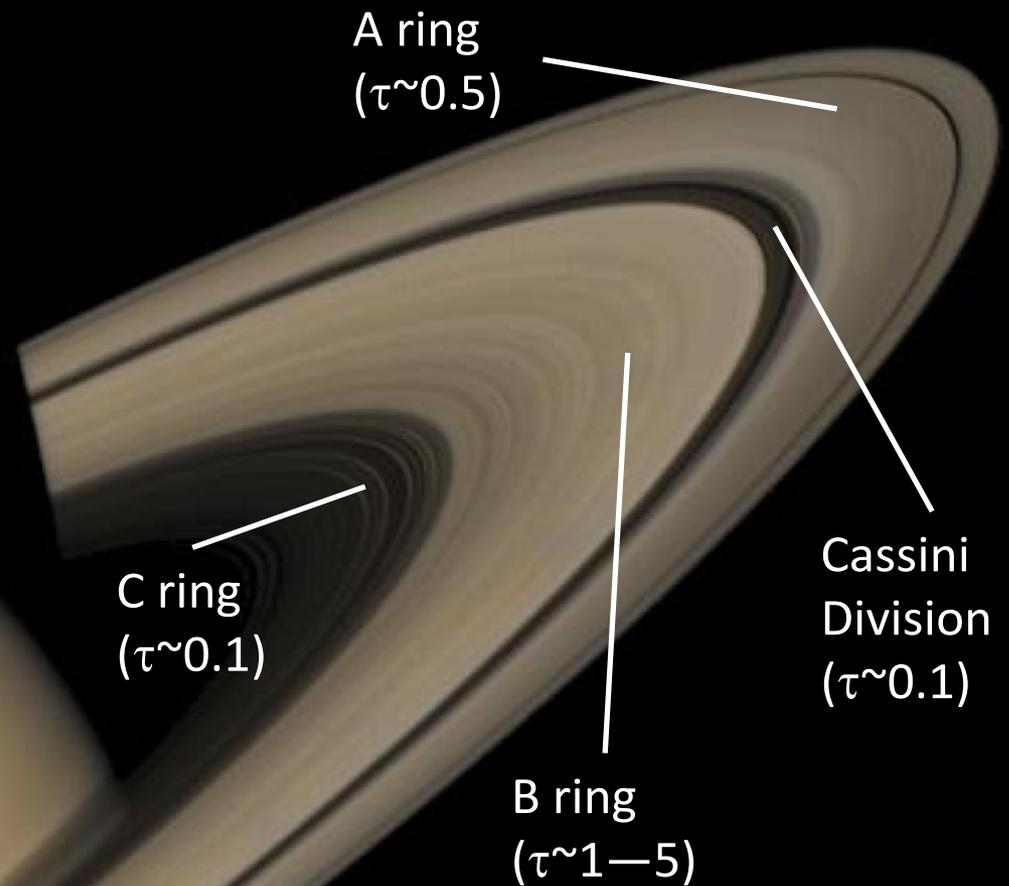
Saturn's rings



Debris Disk HR 4796A
(Thalmann+ 2011)



Saturn's Rings



- Made from Icy Particles
- In visible light: sun light reflected on ring particles
- Optically thick A & B rings much brighter; optically thin C ring & Cassini Division fainter
- How the rings look in thermal emission?

Data

- MIR Camera COMICS on Subaru Telescope with 8.2 m Primary Mirror on Maunakea
- Multi-wavelength Imaging at 8.8, 9.7, 10.5, 11.7, 12.5, 17.7, 18.8, 20.5, 24.5 μm
- Observed in 2008/1/23 by G. Orton, L. Fletcher et al. (unpublished)
- Almost edge-on rings to observer/Sun



Astrometrical parameters

Epoch (UT)	r (au)	Δ (au)	B' (deg)	B (deg)	α (deg)
January 23, 2008	9.27	8.45	-8.7	-7.2	3.5

Sun-Saturn

Solar El.

Phase Ang.

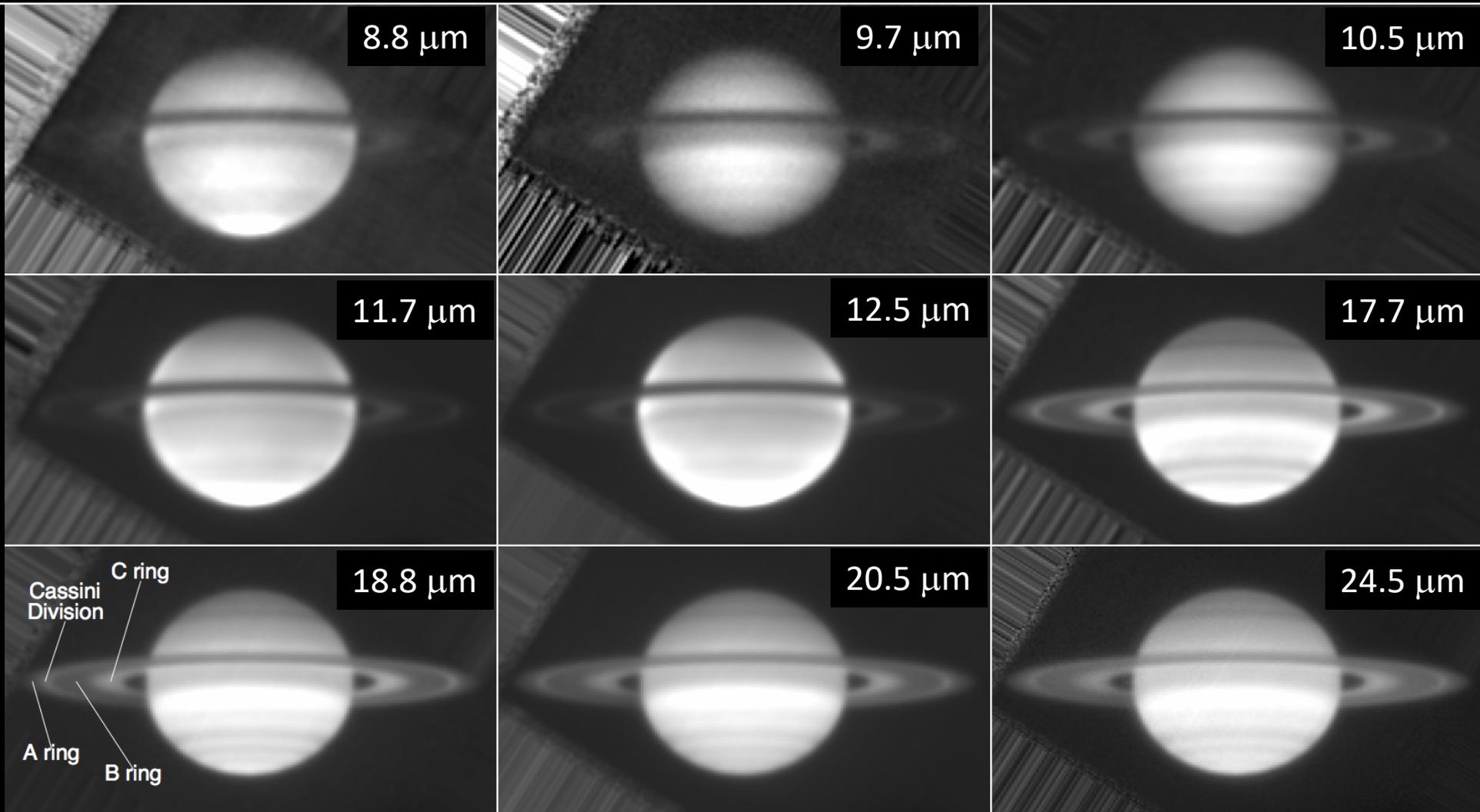
(above ring plane)

Earth-Saturn

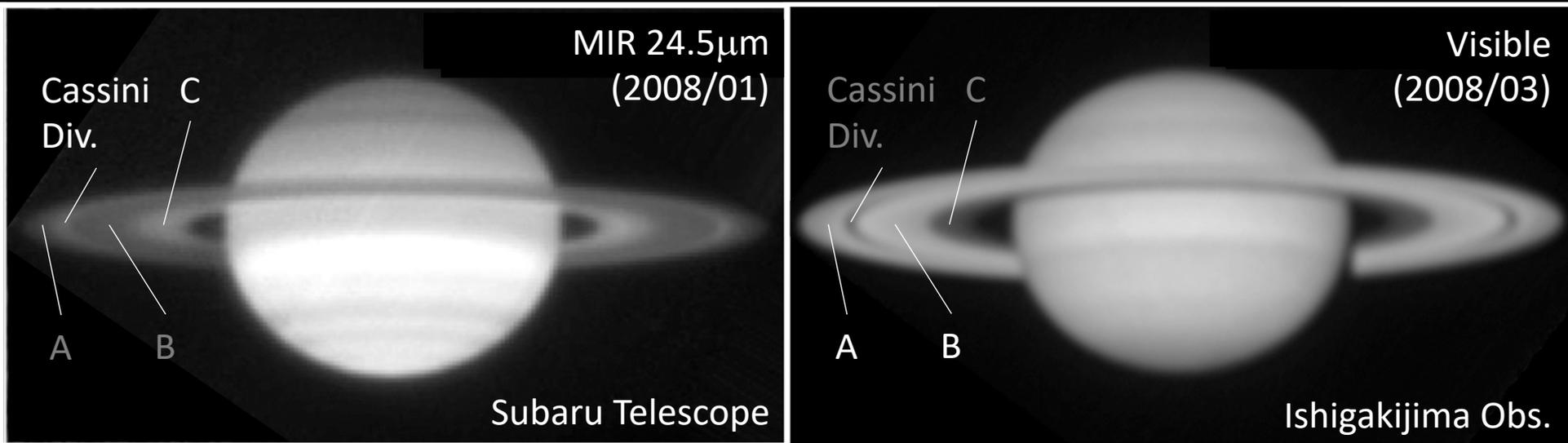
Earth El.

Ring Brightness

Spatial Res: 0.38-0.67" = a few 1000 km
(Highest-ever in MIR from ground)



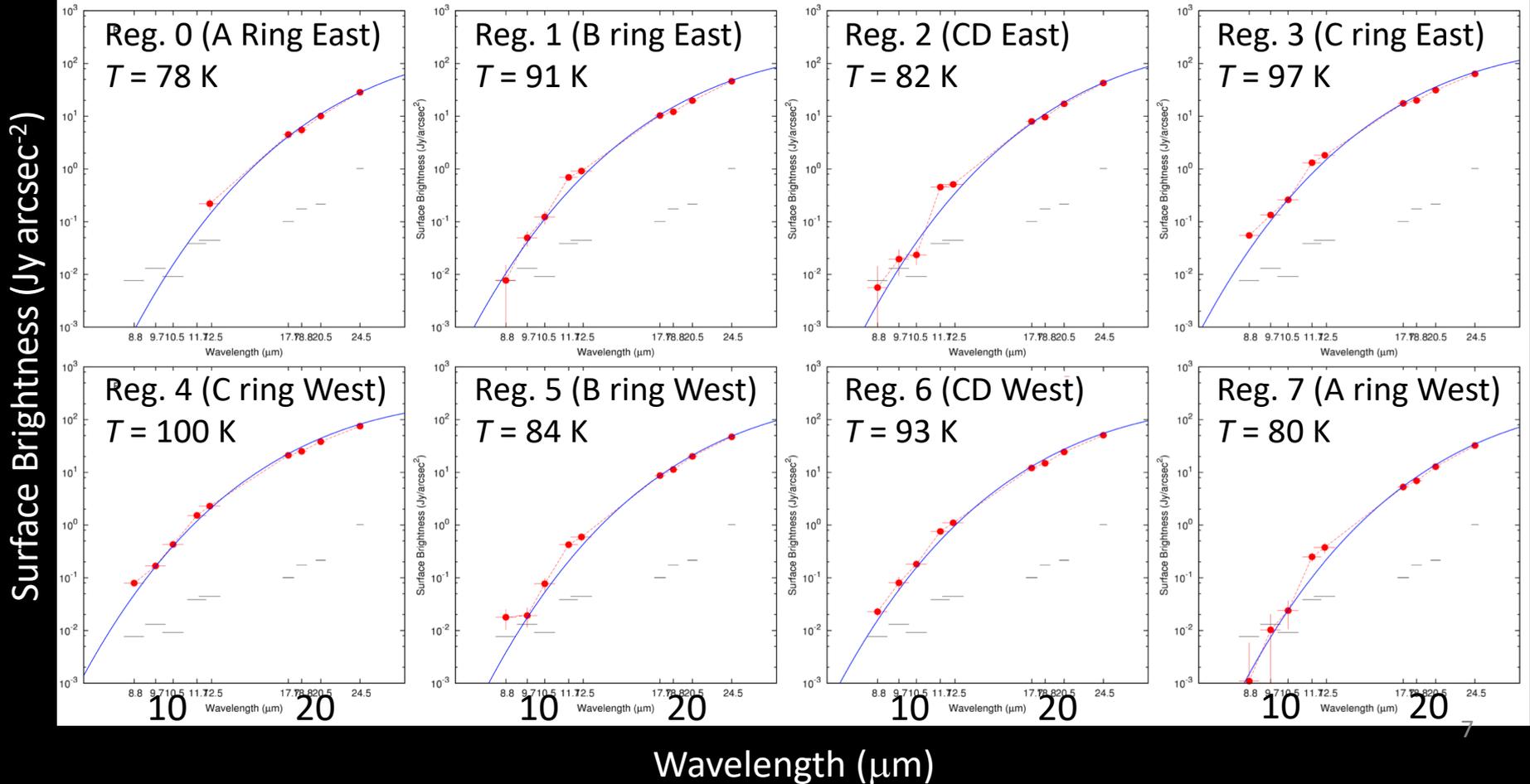
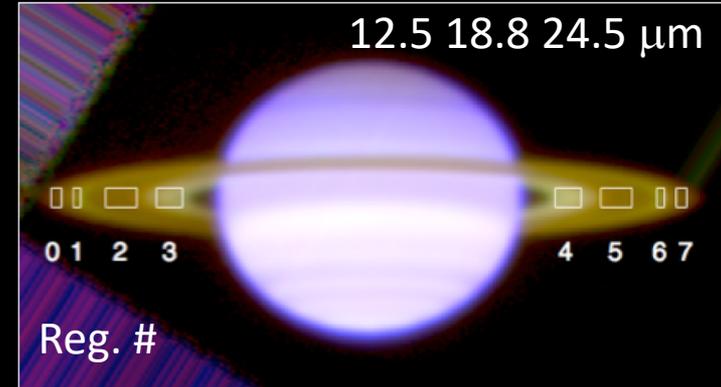
Comparison w/ visible image



- Compared with visible image by Ishigakijima Obs.
- C ring and Cassini Div. are brighter than B and A rings in MIR while C rings and Cassini Div. are always fainter in visible
- Ring brightness contrasts in MIR and visible are reversed

SEDs of Rings

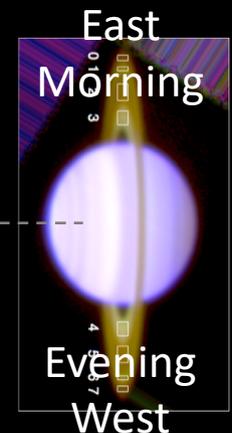
- Blackbody peaking at $> 25 \mu\text{m}$
- No significant dust features



Ring Temperatures

- SED fit w/ BB (assuming τ , measured from stellar occultation)
 - Physical Temperature of Ring Particles
 - Brightness $I_\nu(\lambda) = \beta B_\nu(\lambda, T)$
 - Filling Factor $\beta = 1 - \exp(-\tau/|\sin B|)$

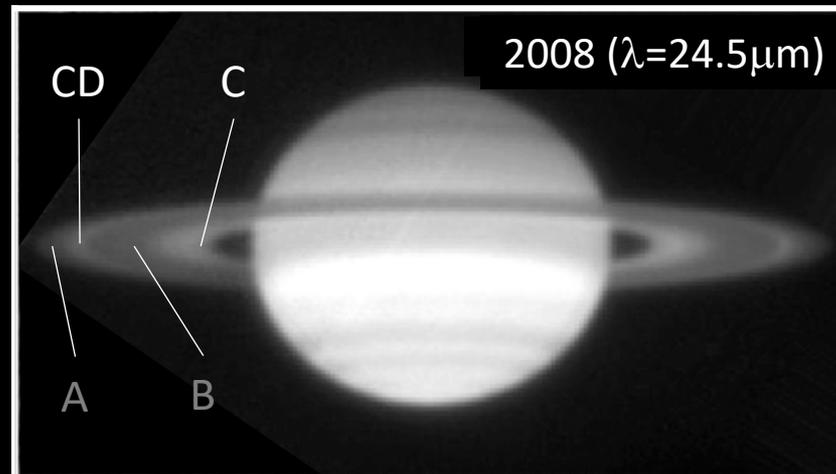
Region	Ring	T by COMICS (K)					
		$\tau = 0.05$	0.1	0.2	0.5	1	2
0	A ring (East)	–	–	80	78	78	–
1	Cassini Division (East)	97	91	87	–	–	–
2	B ring (East)	–	–	–	82	82	82
3	C ring (East)	102	97	92	–	–	–
4	C ring (West)	105	100	95	–	–	–
5	B ring (West)	–	–	–	84	84	84
6	Cassini Division (West)	99	93	89	–	–	–
7	A ring (West)	–	–	82	80	80	–



- C ring & Cassini Div. warmer than B, A rings
- West side warmer than east side in each ring

Discussion

1. In the MIR, optically thin C ring and Cassini Div. were brighter than optically thick A & B rings in 2008
2. Temperatures in C ring and Cassini Div. were higher than A & B rings
3. West side is generally warmer than the east over the rings
 - Accounted for by eclipse cooling in Saturn's shadow (E=morning, W=evening)



- (1) & (2) related to each other
- Surface Brightness in Thermal Emission: $I_{\nu}(\lambda) = \beta B_{\nu}(\lambda, T)$
- Filling factors in C ring and CD were lower than A & B rings
- BUT, effect of higher temperatures in C & CD overcome lower filling factors

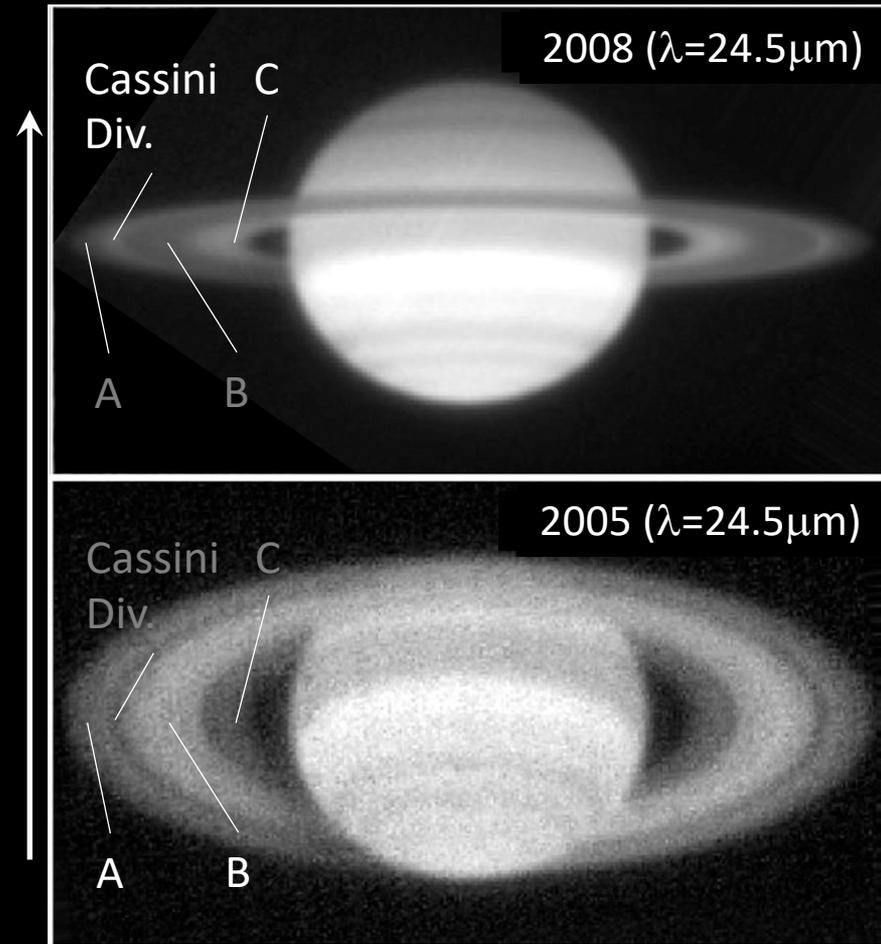
Variation from 2005 to 2008

- Compared with data in 2005 April, when ring opening was larger

Epoch (UT)	r (au)	Δ (au)	B' (deg)	B (deg)	α (deg)
January 23, 2008	9.27	8.45	-8.7	-7.2	3.5
April 30, 2005	9.07	9.33	-21.9	-23.6	6.1

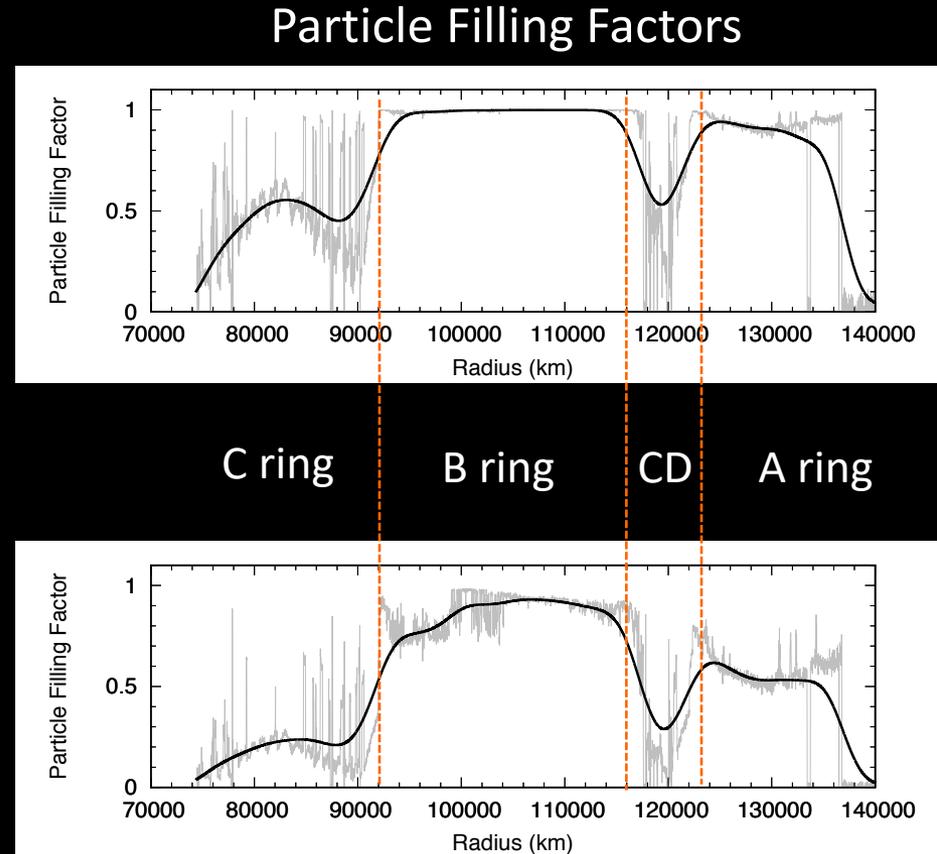
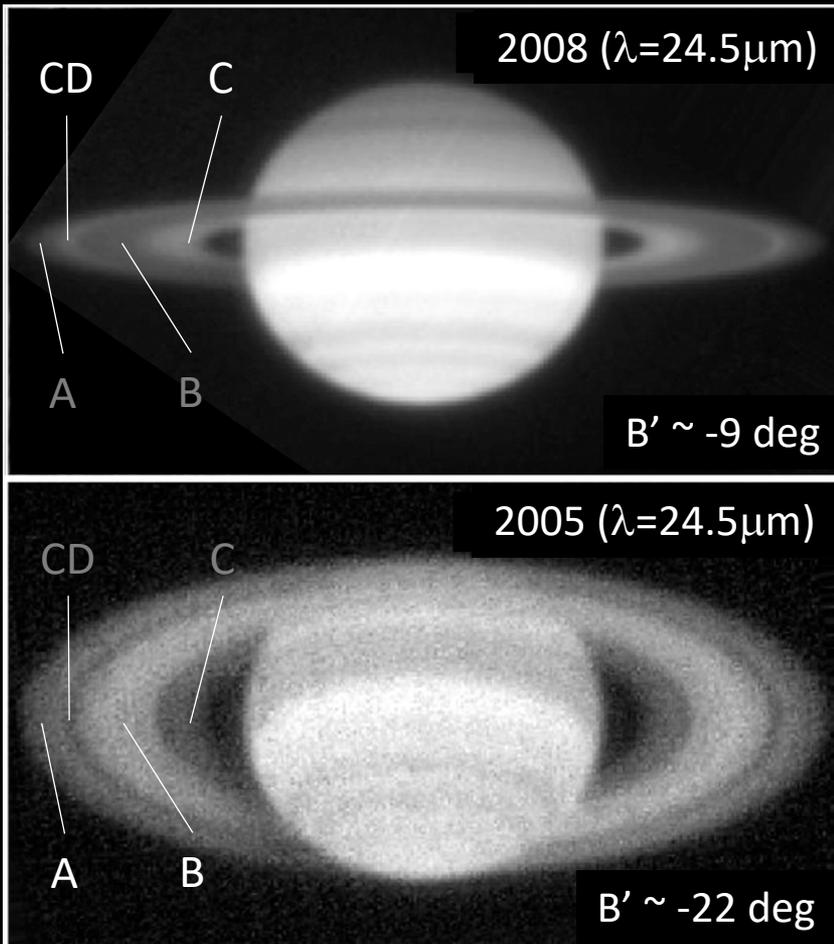
(above ring plane)
Solar El. Earth El.

- Contrast reversed in 2005-2008
- Could be explained by change in filling factors
- Larger opening angle makes larger contrast in optical depths
- C ring & CD fainter



(See Fujiwara et al. 2017 A&A for quantitative discussion)

Temporal Variation in Filling Factors



- Larger ring opening angle makes larger contrast in filling factor

Summery

- MIR image of Saturn from Subaru/COMICS
 - Measurement of Ring Brightness & Temperatures (Highest-ever spatial resolution from the ground)
 - C ring and Cassini Div. were warmer than B and A rings
 - In 2008 C ring and Cassini Div. were brighter than B and A rings (opposite contrast to visible light)
 - MIR Brightness Contrast became inverse from 2005 to 2008
- Observed temporal variation in the MIR brightness contrast is interpreted as a result of a seasonal effect with changing elevations of the Sun and observer above the ring plane